



## Simulations of convective storms in Jupiter with an updated version of a three-dimensional model of moist convection

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Moist convective storms powered by the release of latent heat in rising air parcels are a key element of the meteorology of the Gas Giants [1] and are suspected to play also an important role in the atmospheric dynamics of the Ice Giants [2]. In Jupiter convective storms of different spatial scales occur with different frequencies, from short-lived localized storms [3] to longer-lived storms able to trigger planetary-scale disturbances that develop in cycles of several years [4].

Several models with different approaches have been developed to study moist convection in Jupiter and other planets [5-8]. Three-dimensional cloud resolving models are computationally expensive but have the advantage of allowing the study of the motions generated in the storm and they can also take into account the effects of the three-dimensional Coriolis force in the evolution of the storm. We have used an updated version of a three-dimensional Anelastic Model of Moist Convection [9-11] to explore the development of convective storms in Jupiter. We have improved the dynamical core of the model increasing the stability of the model, which allows us to simulate the dynamics of the development of the storms for longer time ranges than previous simulations presented with this model.

Here we will present results of new simulations of moist convective storms in Jupiter. We simulated the onset and initial development of the storms in a series of different scenarios of condensables abundances to study under which conditions it is possible to trigger convective storms. We tested different abundances of the condensables, relative humidities and fractions of condensates carried by the storm. We pay particular attention to the capacity of the storm to generate convective downdrafts with the potential to desiccate the volatiles of the upper atmosphere [12, 13].

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