



Recycling of Planetary Proto-Atmospheres

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Protoplanets formed by core accretion can become massive enough to accrete gas from the disk they are born in. If the planetary proto-atmosphere exceeds a critical mass, runaway gas accretion starts and the planetary atmosphere collapses into a gas giant. In recent years, many close-in super-Earths have been observed which raises the question on how they avoided becoming hot Jupiters. We investigate the recycling hypothesis as a possible mechanism to avoid the collapse of the atmosphere. We use three-dimensional radiation-hydrodynamics to simulate the formation of proto-atmosphere in the local frame around the planet. In post-processing we use tracer particles to calculate the shape of the atmosphere and determine the non-uniform recycling timescale in a quantitative manner. Our simulations converge to a quasi-steady state where the velocity field of the gas does not change anymore. For the parameter space explored, $a = 0.1 \text{ au}$, $m_c \in [1, 2, 5, 10] M_{\text{Earth}}$, we find that recycling of the atmosphere counteracts the collapse by preventing the gas from cooling efficiently.