



Crust, mantle and core structure of Mars from InSight seismic data

Amir Khan¹ and the the InSight team*

¹Institute of Geophysics, ETH Zurich, Zurich, Switzerland (akhan@ethz.ch)

*A full list of authors appears at the end of the abstract

With the deployment of a seismometer on the surface of Mars as part of NASA's InSight mission, the Seismic Experiment for Interior Structure (SEIS) has been collecting continuous data since early 2019.

The primary goal of InSight is to improve our understanding of the internal structure and dynamics of Mars, in particular crust, mantle, and core. Here we describe constraints on the structure of the mantle of Mars based on inversion of seismic body wave arrivals from a number of low-frequency marsquakes.

We consider 8 of the largest (moment magnitude is estimated to be between 3 and 4) low-frequency events with dominant energy below 1 Hz for which P- and S-waves are identifiable, enabling epicentral distance estimation.

The 8 events occur in the distance range 25-75 degrees. Body wave arrivals that include the main P- and S-waves, surface reflections (PP, PPP, SS, SSS), and core reflections (ScS) are picked using a set of complimentary methods that allows to check for consistency. The resultant set of differential travel times (PP-P, PPP-P, SS-S,...) are subsequently inverted for radial profiles of seismic P- and S-wave velocity, core size and mean density, and epicentral location of the events. To determine interior structure, we rely on independent methods as a means of assessing the robustness of the results.

We present a radial velocity model for the upper mantle of Mars, with implications for the thermo-chemical evolution of the planet that match a cooling, differentiated body, and a thick lithosphere. Based on the location of the events, we are able to constrain structure to the core-mantle-boundary, including the size of the core and its mean density that point to large liquid and relatively light core, implying a significant complement of light alloying elements. Our estimate of the average crustal thickness as seen by all events is compatible with the local crustal

thickness at the InSight landing determined from observations of converted phases.

the InSight team: Simon C. Stähler, Amir Khan, W. Bruce Banerdt, Philippe Lognonné, Domenico Giardini, Savas Ceylan, Mélanie Drilleau, A. Cecilia Duran, Raphaël Garcia, Quancheng Huang, Doyeon Kim, Vedran Lekic, Henri Samuel, Martin Schimmel, Nicholas Schmerr, David Sollberger, Éléonore Stutzmann, Zongbo Xu, Daniele Antonangeli, Constantinos Charalambous, Paul Davis, Jessica C.E. Irving, Taichi Kawamura, Martin Knapmeyer, Ross Maguire, Angela G. Marusiak, Mark P. Panning, Clément Perrin, Ana-Catalina Plesa, Attilio Rivoldini, Cédric Schmelzbach, Géraldine Zenhäusern, Éric Beucler, John Clinton, Nikolaj Dahmen, Martin van Driel, Tamara Gudkova, Anna Horleston, W. Thomas Pike, Matthieu Plasman, Suzanne E. Smrekar, Adrien Broquet, Constantinos Charalambous, Foivos Karakostas, Scott M. McLennan, Chloe Michaut, William T. Pike, Baptiste Pinot, John-Robert Scholz, Rudolf Widmer-Schmidrig, Tilman Spohn, Brigitte Knapmeyer-Endrun, Felix Bissig, Rakshit Joshi, Benoit Tauzin, Saikiran Tharimena, Nicolas Compaire, Ludovic Margerin, Ebru Bozdogan, Mark Wieczorek, Lu Pan, Catherine Johnson, Nienke Brinkman, Anna Mittelholz