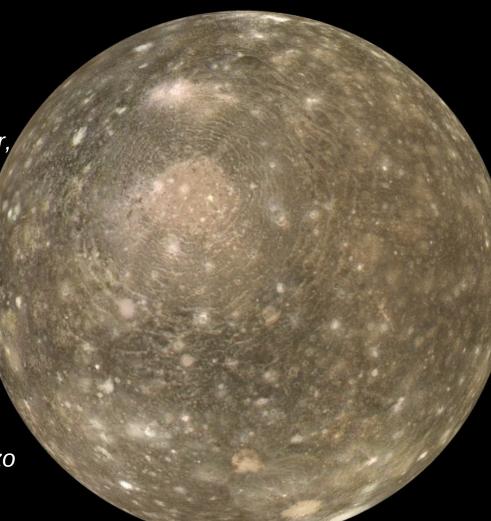
The Particle Environment Package on board JUICE: What Can We Learn about Callisto's Atmosphere and Space Environment?

A. Galli, A. Vorburger, S.R. Carberry Mogan E. Roussos, G. Stenberg-Wieser, P. Wurz, M. Föhn, N. Krupp, S. Barabash, Y. Futaana, P.C. Brandt, P. Kollmann, D. Haggerty, G. Jones, R.E. Johnson, O.J. Tucker, S. Simon, T. Tippens, L. Liuzzo



University of Bern, Switzerland: Center for Space Science, NYU Abu Dhabi, UAE; Max Planck Institute for Solar System Research, Germany; Swedish Institute of Space Physics, Sweden; Johns Hopkins University, USA; University College London, UK; University of Virginia, USA; NASA Goddard Space Flight Center, USA; School of Earth & Atmospheric Sciences, Georgia Institute of Technology, USA; Space Sciences Laboratory, University of California, USA

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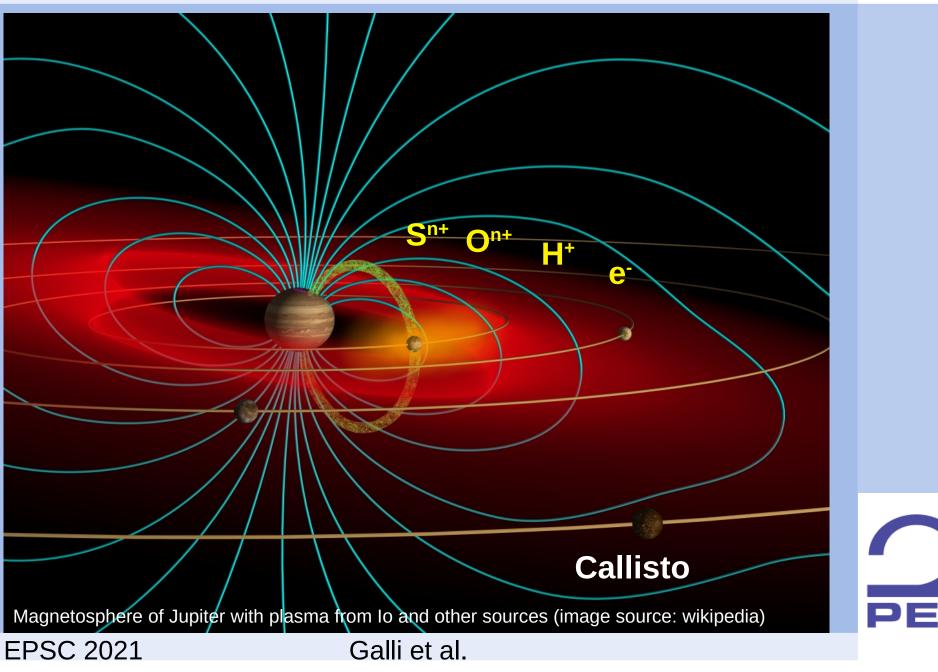
Presentation EPSC2021-169

The Jupiter Icy Moons Explorer (JUICE) will fly by Callisto 21 times before entering into Ganymede's orbit

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JUICE trajectory implications for Callisto observations

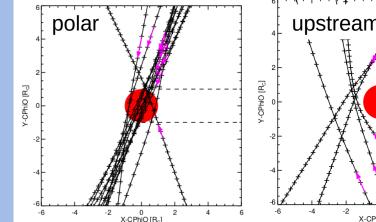
plasma

wake_

[B_c]

(Baseline: CREMA 5.0 for JUICE launch in September 2021, Reference: www.cosmos.esa.int/web/spice/spice-for-juice)

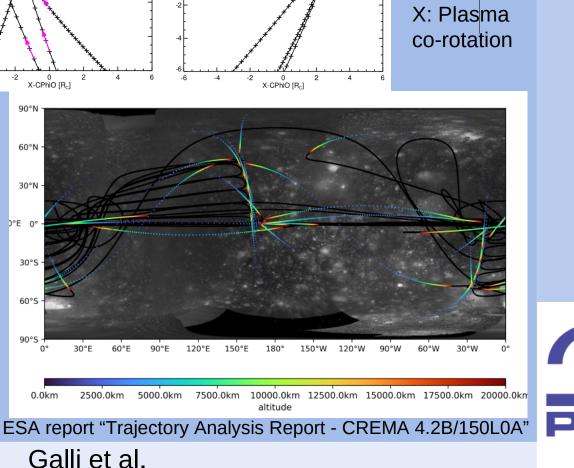
21 Callisto flybys: 12x polar, 6x upstream, 3x downstream



JUICE ground tracks on Callisto's surface: **12 flybys at h < 500 km!**

Polar and equatorial regions covered

Check it out yourself at https://juicept.esac.esa.int



downstream

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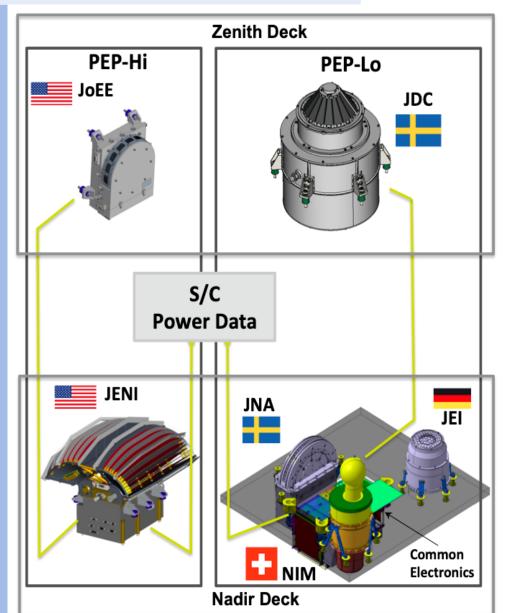
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PEP: Combine in-situ <u>neutrals</u>, <u>ions</u>, <u>electrons</u>, and remote <u>Energetic Neutral Atoms</u> into one big picture

Particle Environment Package (PEP) = six different instruments:

- <u>NIM:</u> Neutral gas and ion mass spectrometer (thermal energies)
- <u>JDC</u>: Ion spectrometer and mass analyzer (electron capabilities)
- <u>JEI:</u> Electron spectrometer (ion capabilities)
- <u>JoEE</u>: Energetic electrons spectrometer
- JNA: Low energy ENA imager
- JENI: Energetic ion spectrometer and ENA imager (electron capabilities)

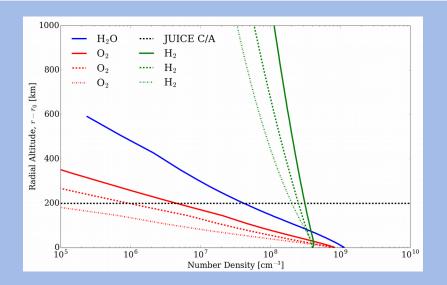


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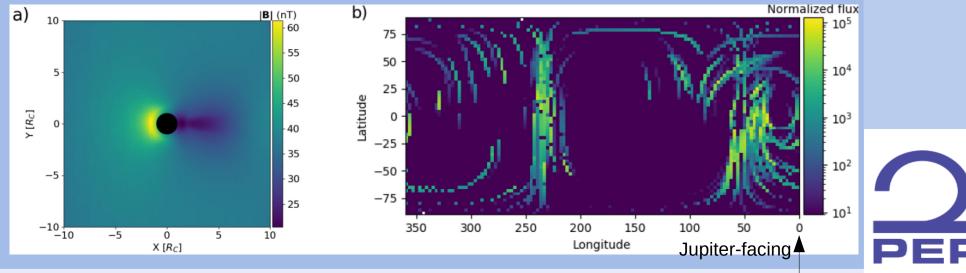
Planning PEP Operations during Callisto flybys: Neutrals, plasma, and Energetic Neutral Atoms



Predicted profiles of the major atmospheric neutral species (Carberry Mogan et al., 2021) for various solar zenith angles; solid lines: 0°, dashed lines: 90°, dotted lines: 180°; JUICE closest approach at 200 km altitude; detection threshold would be 1 cm⁻³

Magnetic field structure (left) and predicted ENA emissions from Callisto (right panel), caused by energetic ions charge-exchanging with neutrals

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Summary

- JUICE trajectory 5.0: 21 flybys in total, 12 flybys at altitude < 500 km, dayside and nightside. Good for Callisto surface coverage; for plasma science and atmospheric loss studies more downstream flybys would be preferable.
- Neutral densities: Low flyby altitude is crucial to detect heavy species, in particular for nightside flybys.
- > PEP: Try to get parallel electron, ion, and neutrals data on the way to closest approach, after closest approach no useful neutrals measurements possible.
- > PEP: ~10s shortest measurement interval except for electron data (~1s).
- > PEP: Background rates due to radiation levels in Jovian magnetosphere at Callisto much lower than near Europa or Ganymede.
- > Implications for models and theory:
 - \rightarrow Ongoing improvement of atmosphere models (1d, 2d, 3d, collisions,...).
 - \rightarrow Combine plasma and atmosphere models of Callisto.
 - \rightarrow First model predictions of Energetic Neutral Atoms images from Callisto done, more details will be published in separate paper.
- > PEP input for models: more accurate and localized measurements of plasma, neutral atmosphere (spatial distribution), and surface composition to better constrain models of the Callisto atmosphere and plasma environment.
- > Questions and comments? Write to andre.galli@unibe.ch. Publication of this study is in preparation.



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