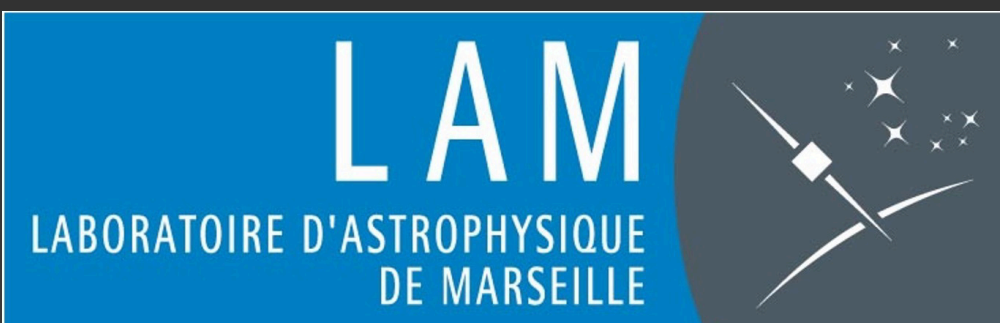


Constraints on the existence of low-mass planets with supercritical hydrospheres

H.G. Vivien¹, A. Aguichine¹, O. Mousis¹, M. Deleuil¹, E. Marcq²

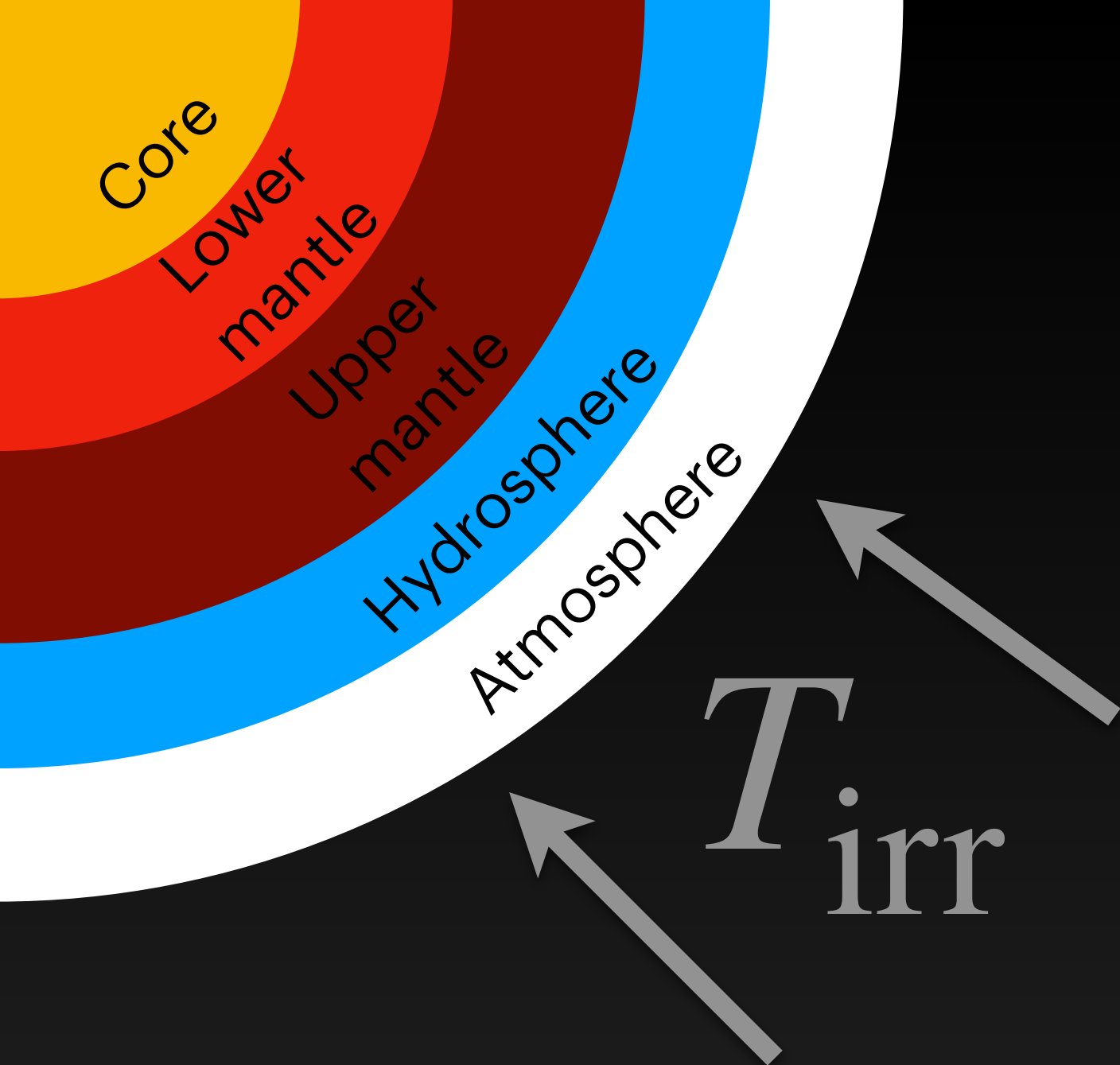


1 - Aix Marseille Univ, CNRS, CNES, LAM, Institut Origines, Marseille, France
2 - LATMOS/IPSL, UVSQ Université Paris-Saclay, Sorbonne Université, CNRS, Guyancourt, France

Context & goals

Existence of Earth-like ocean worlds

- Earth-like mass planets are the epitome of exoplanet detection
- Léger et al. 2004: theoretical existence of water worlds
- Water rich planets could explain the gap between super-Earth and sub-Neptune
- Possible existence of water rich $0.2 - 1M_{\oplus}$ planets
 - Atmosphere stability
 - Core compression



Model

Methods and parameters

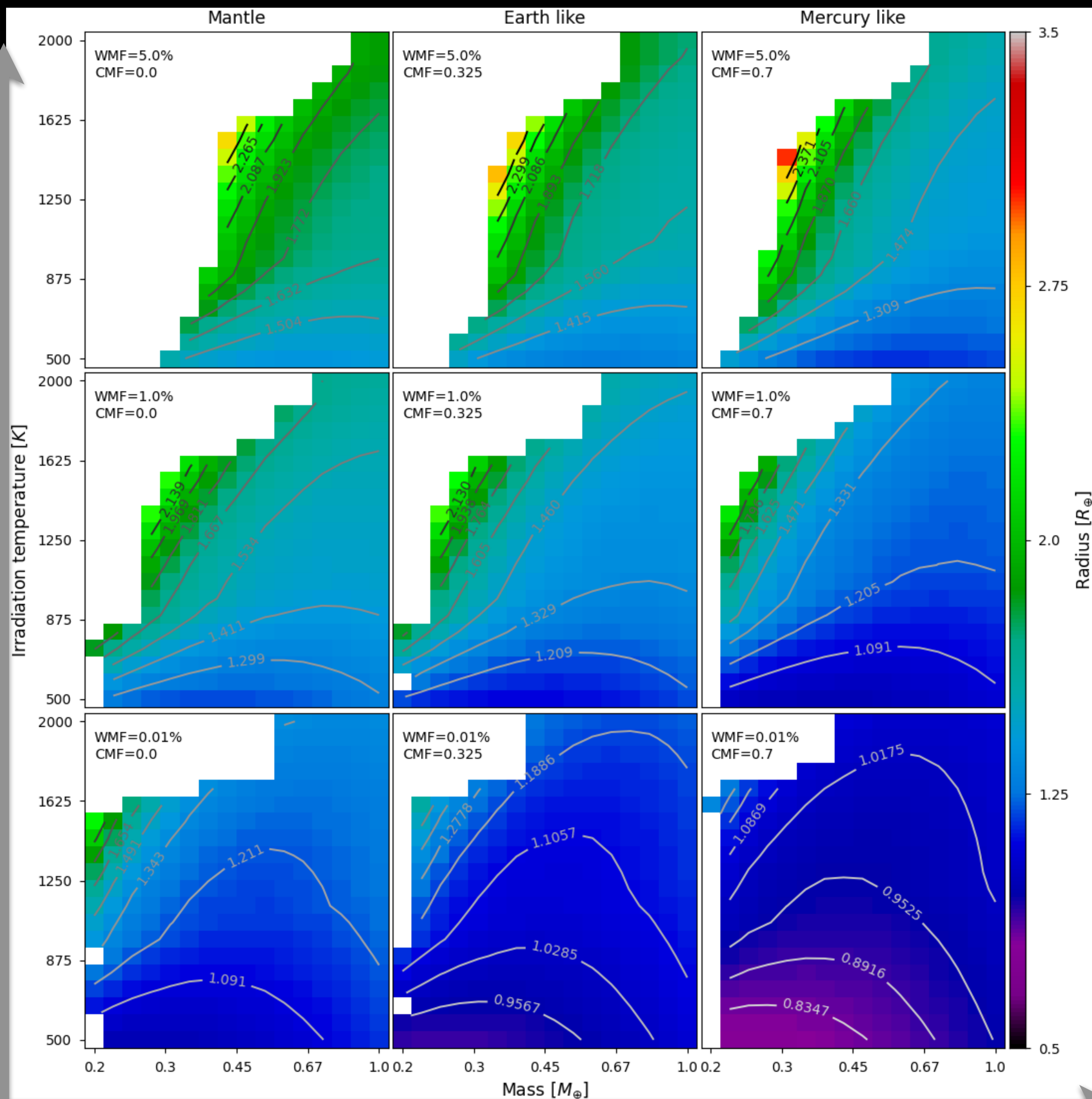
Explored composition:

- CMF: 0, 0.325, 0.7
- WMF: 0.01%, 1%, 5%
- M_p : $0.2 - 1M_{\oplus}$
- T_{irr} : $500 - 2000K$

- Self consistent interior and atmosphere model
 - Implementation of an upgraded convergence scheme
 - The atmosphere grid has been extended to surface gravity as low as $1m \cdot s^{-2}$
 - Hydrostatic stability is assessed for each planet

Brugger et al. 2017, Marcq et al. 2017, 2019,
Mousis et al. 2020, Aguichine et al. 2021

Increasing WMF



Increasing CMF

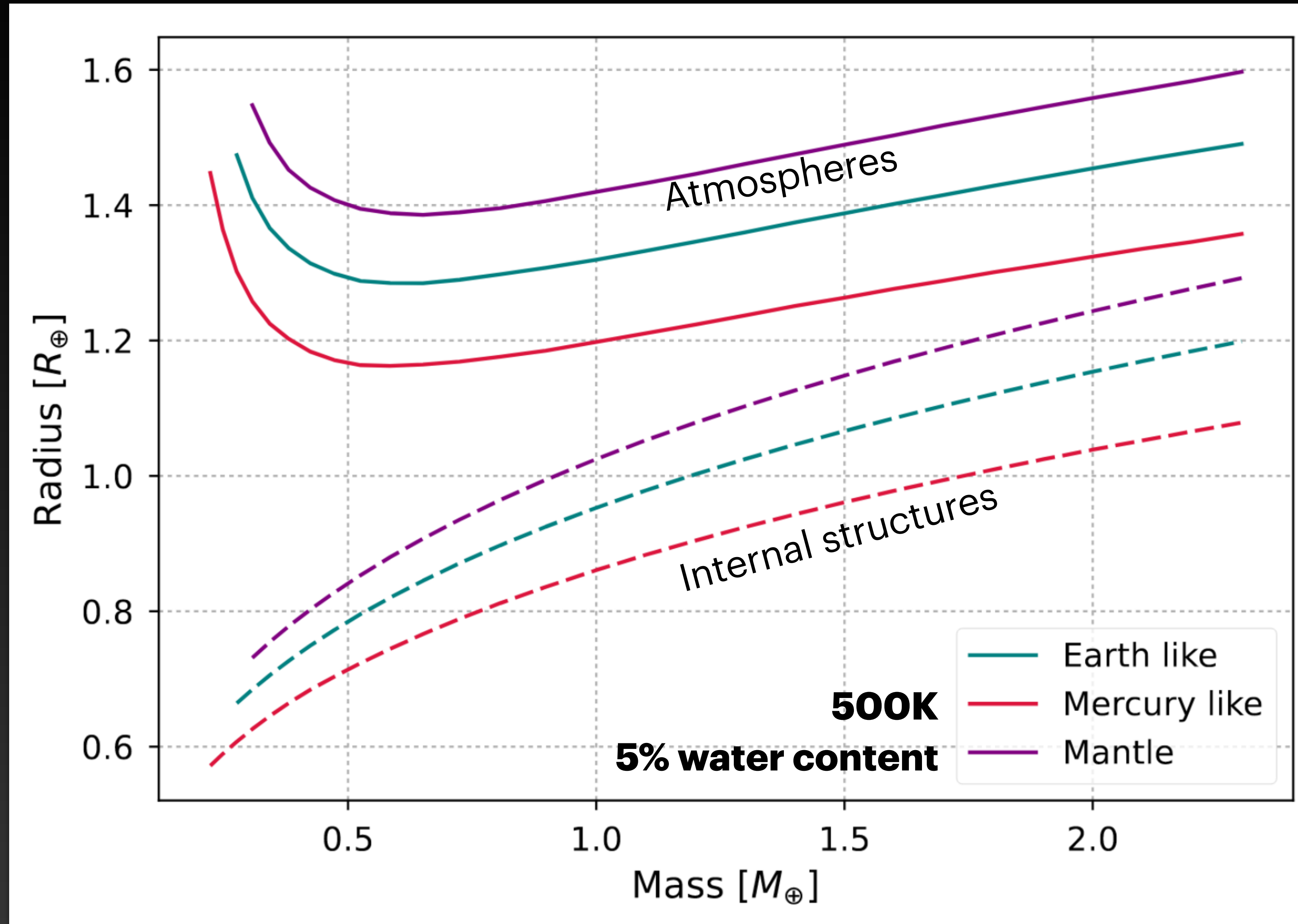
Radius inflation

- The atmosphere boundary is defined at 20mbar
 - White areas → unstable atmosphere
- Atmosphere inflation for low mass and irradiation

Impact of the CMF

→ Radius inflation at low masses

→ Influence of the CMF on the inflation



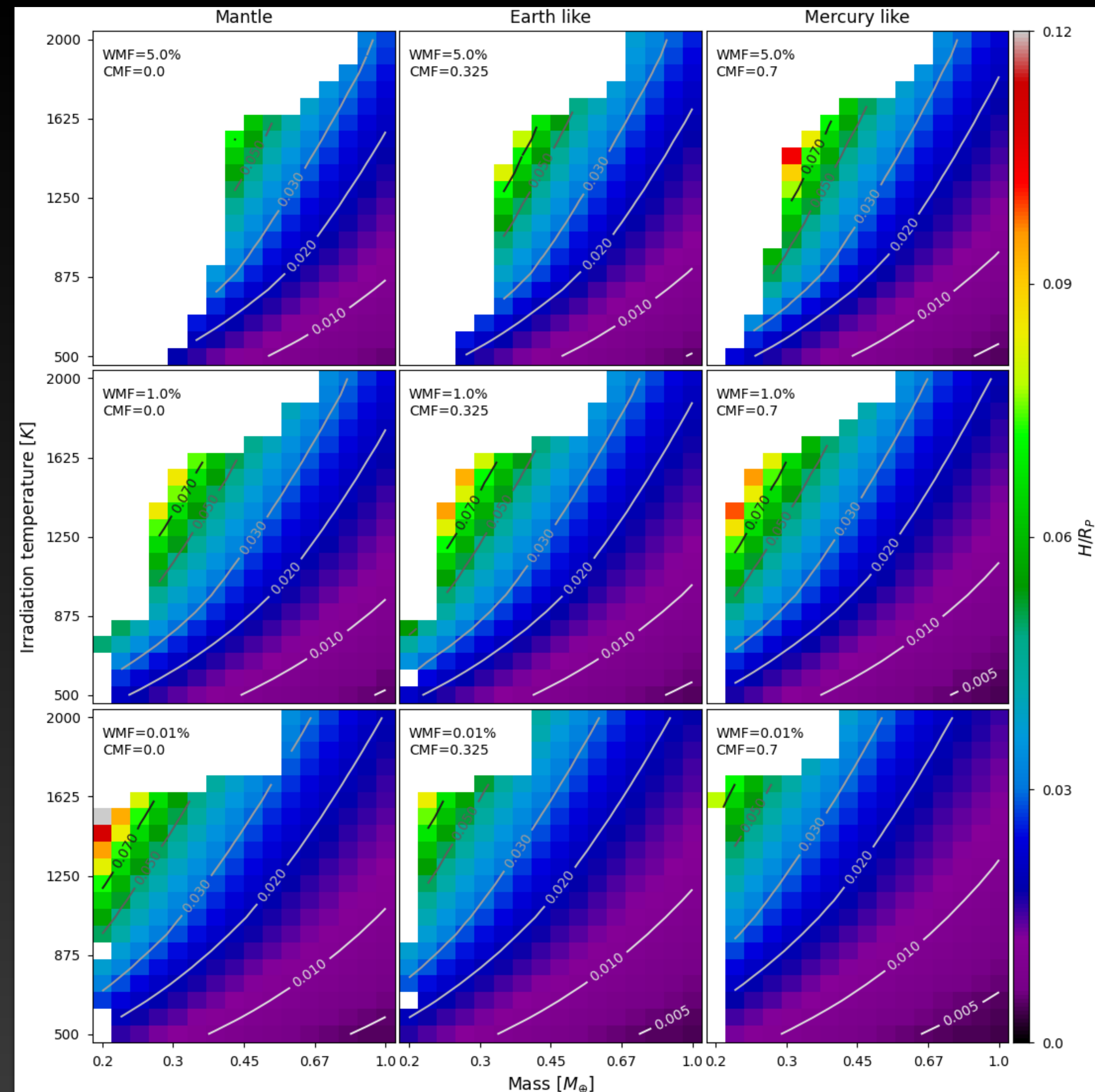
Stability criterion

$$H = \frac{k_B T_{\text{irr}} R_P^2}{GM_P m_{\text{H}_2\text{O}}} \Rightarrow H > \eta R_P$$

→ Atmosphere is lost quickly
when η is greater than ~ 0.1

Similar to $\Lambda = R_P/H$ and
 R_P/R_B for H/He rich planets

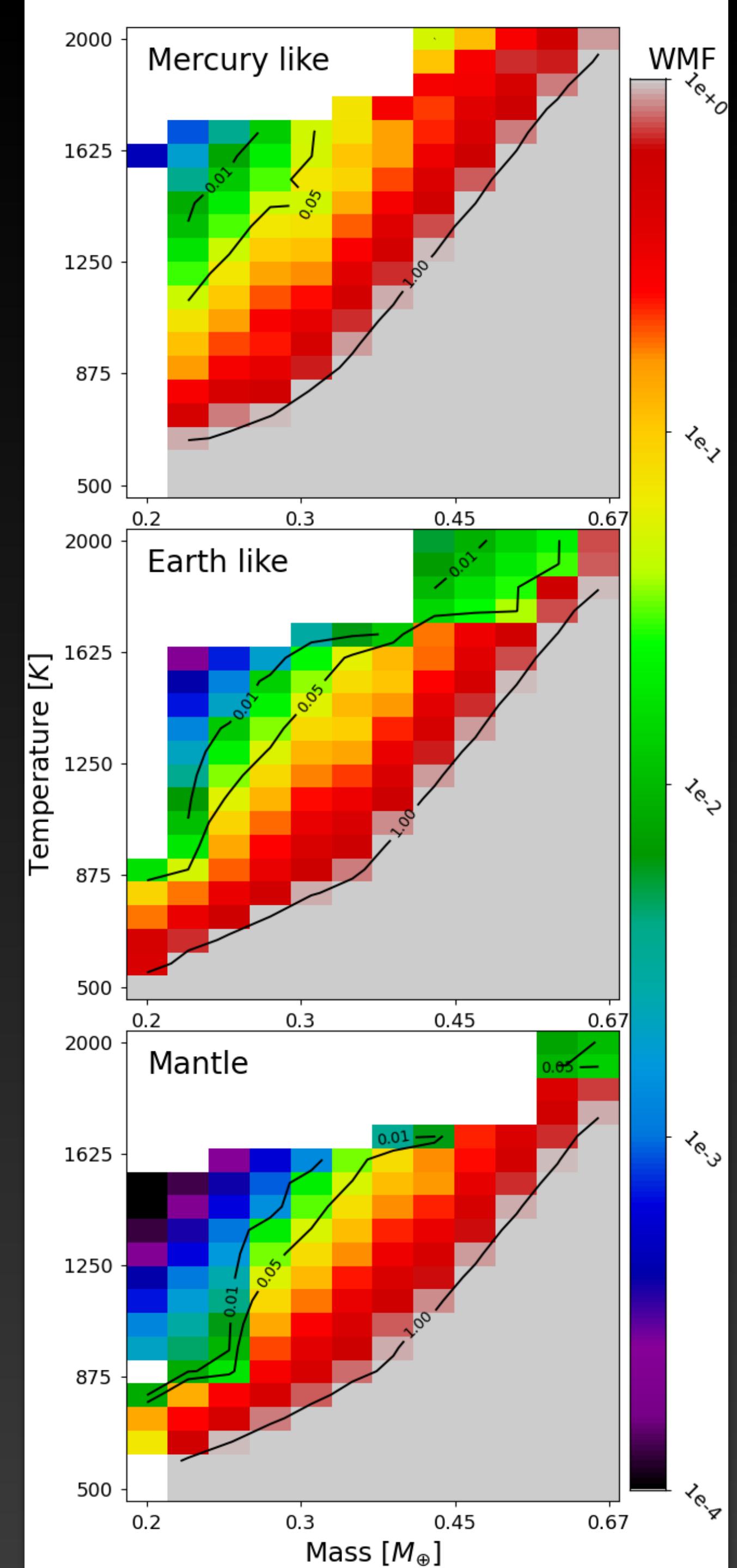
Fossati et al. 2017, Owen & Wu 2016



Theoretical maximum WMF

- For each M_p and T_{irr} we fit R_p as a function of the WMF
- We find the radius for which $\eta = 0.1$

→ Sharp transition between 100% water sustainable and complete evaporation



Keys points

- Model is fully self-consistent, unlike previous studies
- Irradiated low-mass planets can sustain a steam atmosphere under certain conditions
- Internal structure compression plays an important role even at low WMF
- Internal structure is itself noticeably impacted even at low WMF
- Transition of water content sustainability regime is sharp

References

Aguichine, A., Mousis, O., Deleuil, M., et al. 2021, ApJ, 914, 84. doi:10.3847/1538-4357/abfa99

Brugger, B., Mousis, O., Deleuil, M., et al. 2017, ApJ, 850, 93. doi:10.3847/1538-4357/aa965a

Léger, A., Selsis, F., Sotin, C., et al. 2004, Icarus, 169, 499. doi:10.1016/j.icarus.2004.01.001

Fossati, L., Erkaev, N. V., Lammer, H., et al. 2017, A&A, 598, A90.

doi:10.1051/0004-6361/201629716

Marcq, E., Salvador, A., Massol, H., et al. 2017, Journal of Geophysical Research (Planets), 122, 1539. doi:10.1002/2016JE005224

Marcq, E., Baggio, L., Lefèvre, F., et al. 2019, Icarus, 319, 491. doi:10.1016/j.icarus.2018.10.006

Mousis, O., Deleuil, M., Aguichine, A., et al. 2020, ApJL, 896, L22. doi:10.3847/2041-8213/ab9530

Owen, J. E. & Wu, Y. 2016 ApJ, 717, 107. doi:10.3847/0004-637X/817/2/107

Turbet, M., Bolmont, E., Ehrenreich, D., et al. 2020, A&A, 638, A41.

doi:10.1051/0004-6361/201937151

Appendix 1: Comparison with another model

- Models are compatible
- Internal structure compression has an impact even at low WMF

