

## AUTOMATIC PROCESSING OF CHINESE GF-1 WIDE FIELD OF VIEW IMAGES

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### ABSTRACT:

The wide field of view (WFV) imaging instrument carried on the Chinese GF-1 satellite includes four cameras. Each camera has 200km swath-width that can acquire earth image at the same time and the observation can be repeated within only 4 days. This enables the applications of remote sensing imagery to advance from non-scheduled land-observation to periodically land-monitoring in the areas that use the images in such resolutions. This paper introduces an automatic data analysing and processing technique for the wide-swath images acquired by GF-1 satellite. Firstly, the images are validated by a self-adaptive Gaussian mixture model based cloud detection method to confirm whether they are qualified and suitable to be involved into the automatic processing workflow. Then the ground control points (GCPs) are quickly and automatically matched from the public geo-information products such as the rectified panchromatic images of Landsat-8. Before the geometric correction, the cloud detection results are also used to eliminate the invalid GCPs distributed in the cloud covered areas, which obviously reduces the ratio of blunders of GCPs. The geometric correction module not only rectifies the rational function models (RFMs), but also provides the self-calibration model and parameters for the non-linear distortion, and it is iteratively processed to detect blunders. The maximum geometric distortion in WFV image decreases from about 10-15 pixels to 1-2 pixels when compensated by self-calibration model. The processing experiments involve hundreds of WFV images of GF-1 satellite acquired from June to September 2013, which covers the whole mainland of China. All the processing work can be finished by one operator within 2 days on a desktop computer made up by a second-generation Intel Core-i7 CPU and a 4-solid-State-Disk array. The digital ortho maps (DOM) are automatically generated with 3 arc second Shuttle Radar Topography Mission (SRTM). The geometric accuracies of the generated DOM are 20m for camera-2 and 3, and 30m accuracy for camera-1 and 4. These products are now widely used in the fields of land and resource investigation, environment protection, and agricultural research.

### 1. INTRODUCTION

There are three trends in the development of modern remote sensing technologies. (Li, 2003) The first one is to improve the ground resolution so that users can distinguish more details of the ground covers. The newly launched Worldview-3 satellite which is available to acquire 0.31m panchromatic images reflects this trend. The second one is to improve the spectral resolution so that people can more accurately measure and analyse the chemical composition of land covers. The third one is to shorten the revisit time so that periodically land-monitoring can be achieved. Shorter revisit period can make the remote sensing images very useful in agricultural and environmental monitoring applications.

In order to shorten the revisit time of the widely used pushbroom remote sensing satellite, two methods can be employed. One is to improve the rotation ability of the satellite; the other is to improve the swath width of sensors. Using the former method, users can acquire high ground resolution images at a few important locations. However, the latter method has to be used while acquiring images in short revisit time on large-scale area.

The first satellite of the high resolution land-observation system of China (GF-1 satellite), launched on April 26th, 2013, applied the wide swath width method and thus greatly shortened its revisit time. It carries two imaging system: Panchromatic and Multi-Spectral (PMS) system which is used for common high-resolution observation, and wide field of view (WFV) system which is used for large-scale observation. Figure 1 shows four images that the WFV system observed at the same time. Table 1

lists part of the characteristics of the GF-1 satellite and its imaging systems. As is illustrated in figure 2, the WFV imaging system is made up by four cameras, and each of them has about 16 degrees of viewing angle. The WFV images are most commonly delivered in sensor-corrected (SC) data type, i.e. with square images, RFMs, (Fraser *et al.*, 2006; Zhang *et al.*, 2012) and other metadata. The images of four cameras are delivered separately and they are correspondingly processed separately in this paper.

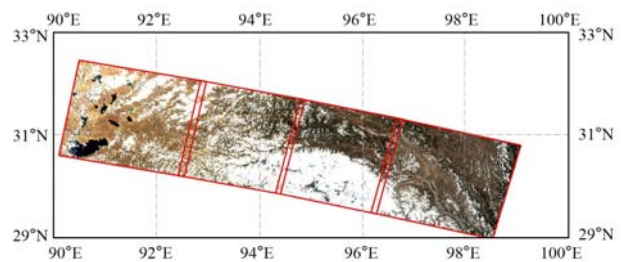


Figure 1. Four SC images that the WFV imaging system observed at the same time.

In order to improve the efficiency of the WFV SC image processing, the research team at Wuhan University developed an automatic processing system. The system can automatically correct the geometric distortions, refine the RFMs by block adjustment, and generate DOMs that have better absolute and relative accuracy than the initial SC images. The automatic processing workflow is described in section 2. The technology of geometric distortion correction is introduced in section 3. In

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section 4, block adjustment experiment using 86 WFV images covering Tibet, China, and generation of DOMs covering the most part of mainland of China are introduced. Section 5 gives the conclusions.

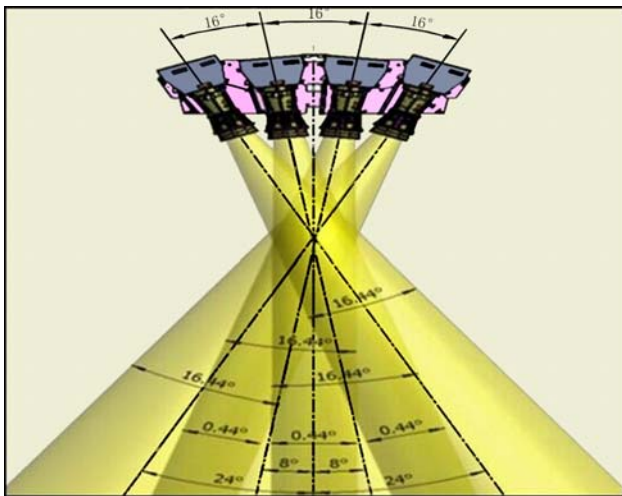


Figure 2. Configuration of WFV imaging system

Items	Parameters	
Orbit type	sun synchronous orbit	
Average Orbit height	644.5km	
descending node	10:15 am local solar time	
<b>Imaging system</b>	<b>PMS</b>	<b>WFV</b>
Revisit time	41 days, 4 days if rotation	4 days
Ground resolution of panchromatic camera	2m at nadir	-
Ground resolution of multispectral camera	8m at nadir	16m at nadir
Swath width	>60km	>800km
Spectral bands ( $\mu m$ )	Pan: 0.45~0.90; B1: 0.45-0.52 B2: 0.52-0.59 B3: 0.63-0.69 B4: 0.77-0.89	B1: 0.45-0.52 B2: 0.52-0.59 B3: 0.63-0.69 B4: 0.77-0.89

Table 1. Part of the parameters of GF-1 satellite and its imaging systems

## 2. WORKFLOW OF AUTOMATIC PROCESSING

The processing workflow of WFV SC images is data driven and is called by the main processing system. The main processing system supplies the automatic processing system with WFV SC images, referencing images (most of them are Landsat-8 panchromatic orthophoto maps), and distortion correction models. The whole workflow is illustrated in figure 3 and can be divided into four main steps, i.e. data preparation, geometric distortion correction, RFM refining and DOM generating.

### 2.1 Data preparation

In the data preparation step, the radiometric correction is optional and for users having special requirement, it can be ignored. Cloud and water detection is used for removing the images that have too much cloud-covered area or water-covered area out of the automatic processing workflow. The removed images can be processed with manual intervention. The detection results (mask images) are also used for eliminating the

mismatched GCPs and tie-points which fall in the cloud or water covered area.

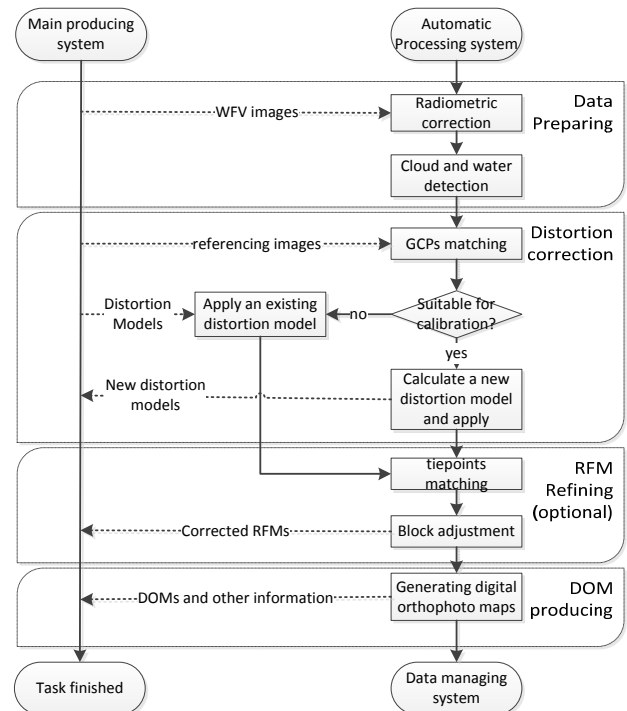
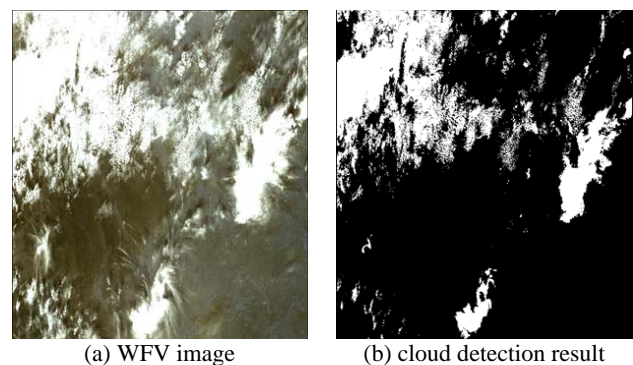


Figure 3. Workflow of the automatic processing system.

A Gaussian mixture model based method is used for cloud detection. Firstly, the initial model parameters are obtained by the intensity histogram of the grey level image which is composed by the corresponding four bands WFV image. Then the parameters are iteratively adjusted with expectation maximization criterion. When the optimal fitting result is reached, the intensity threshold is automatically obtained. Experiments show that when WFV images are processed on a desktop computer having Intel Core-i7 CPU and solid state disk (SSD), the time of obtaining the threshold and generating the mask image are less than 10 seconds and 30 seconds, respectively.

Figure 4 shows one of the cloud-detection results. Most of the thick clouds are successfully detected. However, the detection of thin clouds is not quite good yet. For geometric correction, this result is enough for mismatch elimination, but for image-fusion, other methods should be utilized. (Panem *et al.*, 2005; Wu *et al.*, 2013; Zhu and Woodcock, 2012)



(a) WFV image (b) cloud detection result  
Figure 4. An WFV SC image and the result of the Gaussian mixture model based cloud detection.

## 2.2 Distortion correction

Through the experiments on a large number of WFV images, we have found that there exist non-linear geometric distortions in the WFV SC images. The distortions in different WFV cameras are slightly different, but they are stable in a relatively short period, for example one month. Thus, a well calculated distortion correction model can be used on the images acquired in similar times by the same camera.

In this step, the matched GCPs are used for calculating the distortion correction models and for refining the RFMs. The distribution of GCPs (the mismatches in the cloud or water covered area are already eliminated by the mask images) have to be evaluated to determine whether the image is suitable for calculating the distortion correction model. For an image that is not suitable for calculating a new model, a calibrated model of the same camera image is used.

If a block adjustment is needed, the distortion correction models are used for correcting the image-space position of the GCPs and tie-points, and the already refined RFMs are used to guide the matching of tie-points. The detailed theory and method of distortion correction will be discussed in section 3.

## 2.3 RFM Refinement

Although the RFMs are refined by GCPs in the above step, a block adjustment using both GCPs and tie-points is still important to ensure the relative precision of DOMs (Teo *et al.*, 2010; Toutin, 2003; Toutin, 2006; Xiong and Zhang, 2011; Zhang and Zheng, 2012).

The tie-points are matched with the guidance of GCP-refined RFMs and the constraint of digital elevation model (DEM) in three steps. (Xiong *et al.*, 2012; Xiong and Zhang, 2013) Firstly, the RFM and DEM data are used to compute the approximate topographic relief within the image coverage. Based on the topographic relief, the approximate epipolar lines are constructed and their accuracy is analysed. Secondly, the point matching is executed within the error range of the epipolar line. Finally, the Multi-photo geometrically constrained matching algorithm (Zhang and Gruen, 2006) is adopted to refine the matching result.

After the preparation of GCPs and tie-points, the elevation constrained block adjustment is conducted. The image-space affine transformation is used as adjustable function and after block adjustment, the affine transformation models of all the images are combined into the new RFM parameters.

## 2.4 DOM Generation

Once the distortion correction models and refined RFMs of the WFV images are prepared, the DOMs can be generated. Parallel processing strategy is used to speed up the computation efficiency of DOM generation.

# 3. GEOMETRIC DISTORTION CORRECTION

## 3.1 Error analysis

The geo-referencing errors of pushbroom satellite images are mainly caused by the interior orientation error of the imaging system and the exterior error of the satellite platform.

When RFM is used to transform a point from ground to a pushbroom satellite image, the transforming error (denoted as  $\Delta x$  and  $\Delta y$ ), caused by the exterior orientation error, can be parameterized by the image-space affine transform model. To simplify the using of the affine model, the nominal image-space position  $x_{nom}$  and  $y_{nom}$  (the position just calculated by the RFM and the ground position) are mostly used as arguments, as is denoted:

$$\begin{aligned}\Delta x(x_{nom}, y_{nom}) &= a_0 + a_1 x_{nom} + a_2 y_{nom} \\ \Delta y(x_{nom}, y_{nom}) &= b_0 + b_1 x_{nom} + b_2 y_{nom}\end{aligned}\quad (1)$$

The interior orientation error is mainly caused by lens distortion and errors in the focal length, principal point position, and camera assembly angle. (Jiang *et al.*, 2013; Pan *et al.*, 2013; Zhang *et al.*, 2014) The value and direction of interior orientation error are only related to the real position of CCD pixel, i.e. the real position on the image (denoted as  $x_{mea}$  and  $y_{mea}$ ). However, for pushbroom images, y-coordinate represents the number of scanning lines. Thus, only the x-coordinates of the real image position can be used as the argument in interior orientation error model. In fact, the disparity between the real and nominal image position is relatively small (usually several pixels), which means that the interior and exterior orientation error models have very strong correlation. Thus, these two models cannot be solved separately. To solve this problem, the linear part of interior orientation error is usually combined into the exterior orientation error model, and only the nonlinear part (distortion correction model, denoted as  $D_x$  and  $D_y$ ) is solved separately.

In conclusion, for a ground point, the relationship between its nominal image position and its real image position (measured by image matching) is:

$$\begin{aligned}x_{mea} + D_x(x_{mea}) + \varepsilon_x &= x_{nom} + \Delta x(x_{nom}, y_{nom}) \\ y_{mea} + D_y(x_{mea}) + \varepsilon_y &= y_{nom} + \Delta y(x_{nom}, y_{nom})\end{aligned}\quad (2)$$

where  $x_{mea}$  and  $y_{mea}$  are the real image position,  $\varepsilon_x$  and  $\varepsilon_y$  are the random error of image matching.

## 3.2 Distortion correction model

For the users of GF-1 SC image, the detailed structure of GF-1 WFV camera is hidden in the RFMs, so geometric distortion can only be corrected by an empirical model. When ignoring the geometric distortion ( $D_x$  and  $D_y$ ) in the least squared solving of the affine model ( $\Delta x$  and  $\Delta y$ ), the distribution of image-space residual errors ( $\varepsilon_x$  and  $\varepsilon_y$ ) on the real image-space x-coordinates of the WFV camera is illustrated in figure 5(a).

The residual error illustrated in figure 5 is the combination of the nonlinear geometric distortion and random matching errors. The distortion in x-direction is much more complicated than that in y-direction and is too complicated to be fitted by polynomials (the order maybe every high). However, they all have continuity. Thus, the three-order B-spline function that has six inner nodes is used to fit the geometric distortion curves of the WFV images. One of the fitting results is illustrated in figure 5(b).

### 3.3 Solving strategy

Because of the strong correlation, the distortion correction model and the affine model cannot be solved simultaneously. Furthermore, the fitting of B-spline curves needs uniformly distributed GCPs, and only the images satisfying this condition can be used to calculate their distortion model parameters. To judge whether an image is suitable for calculation, its GCPs are divided into several groups (12 groups for WFV image) based on their image-space x-coordinates (see figure 6). If the number of GCPs in each group is larger than a certain threshold (>20 for WFV image), the image is treated as suitable for calibration.

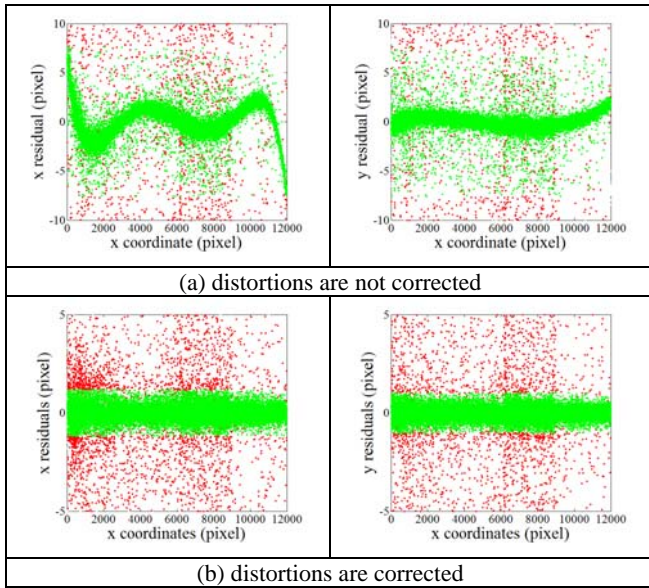


Figure 5. The distribution of the image-space residuals against x-coordinates. The red dots represent the GCPs that are eliminated as outliers, while the green dots are inliers.

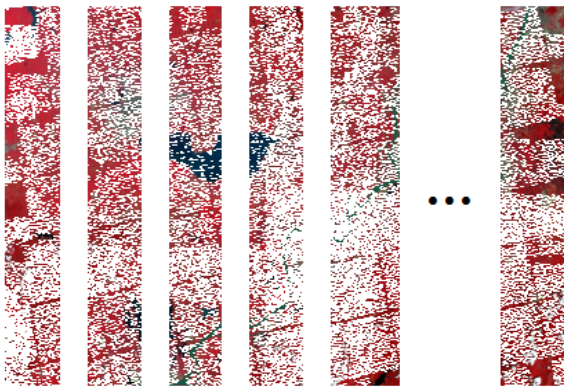


Figure 6. The GCPs are divided into several groups based on the x-coordinates. A white dot represents a GCP.

For the images suitable for calculation, their affine model ( $\Delta x$  and  $\Delta y$ ) are firstly solved by ignoring their distortion. Then, with the already solved affine model, the distortion correction model is solved by the robust least squares method with the loss function:

$$\rho(s) = \log(1 + s) \quad (3)$$

where  $s$  is the value of cost function.

For the images not suitable for calculation, their distortion models should be searched from the data management system. The searched model should be calculated using the image acquired by the same camera, and their shooting time intervals should be less than a month. Then, the searched models are used to correct the image-space position of GCPs. Finally, the corrected GCPs are used to solve the affine models. A RANSAC based method is used to eliminate outliers.

## 4. EXPERIMENTS AND DISCUSSIONS

### 4.1 Block adjustment experiment

This test is to compare the relative precision of DOMs which are generated in two different conditions: (1) the distortion correction models is not used, but the RFMs are block adjusted, (2) the distortion correction model is used and the RFMs are block adjusted.

In this experiment, 86 GF-1 WFV images covering Tibet, China are involved in. All the images were acquired between June, 2013 and September, 2013. The distribution of their DOMs is shown in figure 7. The SRTM data with 90m grid size is used as elevation constraint. The image positions of tie-points are corrected before the block adjustment.

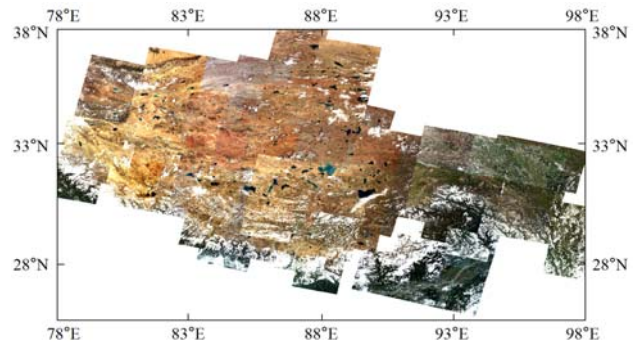


Figure 7. The DOMs of the involved 86 GF-1 WFV images

In condition (2), the image positions of the tie-points are corrected with the distortion correction models, while in condition (1) they are not corrected. The root of mean squared error (RMSE) of the image residuals and the elevation residuals (the tie-points are constrained in elevation) in the block adjustment are listed in table 2. The statistic information shows that the RMSE of image-space residuals and elevation is remarkably reduced if the image-space coordinates are corrected by the distortion correction models.

Items		Condition (1)	Condition (2)
Tie-point number		123977	123977
Detected outliers		266	213
RMSE	Sample (pixel)	1.02	0.35
	Line (pixel)	0.50	0.46
	Elevation (meter)	77.35	22.61
MAX	Sample (pixel)	6.69	3.78
	Line (pixel)	4.31	4.25
	Elevation (meter)	525.02	164.69

Table 2. Statistic information of the block adjustment results of condition (1) and condition (2).

The relative precision in the overlapping area of an image-pair is evaluated as follows: firstly, collect the corresponding points

(from the tie-points list) of the image-pair; secondly, compute the two ground positions with the two RFMs and the SRTM DEM; finally, the mean value and RMSE of the differences in ground positions are calculated. Totally 358 image-pairs are evaluated in this test area. The histograms of the mean value and RMSE of the ground position differences of the image-pairs are illustrated in figure 8. The mean value reflects the overall relative precision of an image-pair. The RMSE reflects the relative precision in parts of the overlapping area and is influenced by the elevation accuracy of SRTM data and the viewing angle of the image-pair.

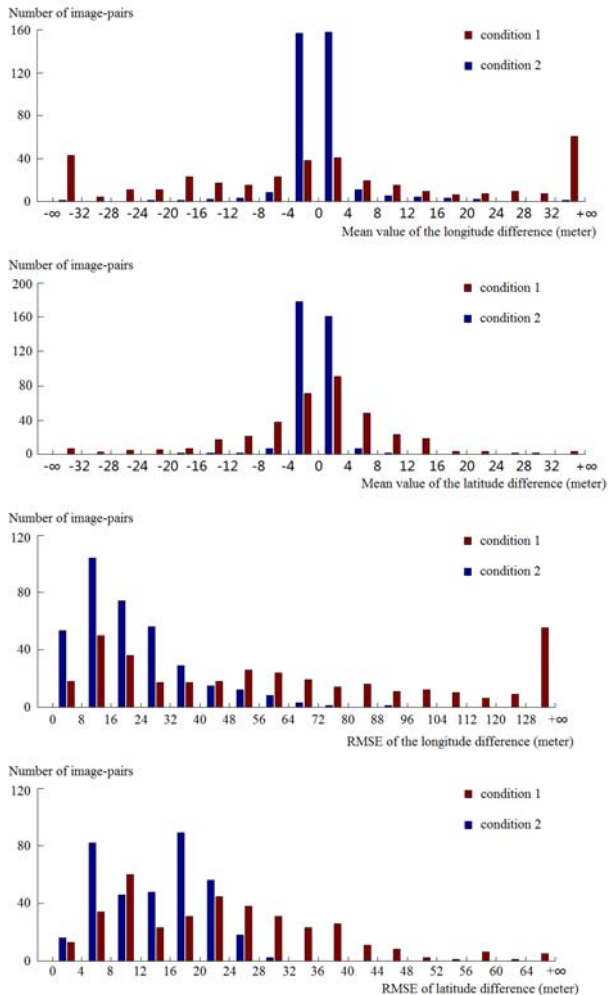


Figure 8. The histograms of the mean value and RMSE of the ground position difference of the corresponding points in all the image-pairs. The relative precision of the whole test area in condition 2 is obviously better than that in condition 1, whether judged by the mean value or by the RMSE.

The results show that the geometric distortion in the GF-1 WFV images strongly influences the relative precision of the DOM products. However, after the correction of geometric distortion, the relative precision is significantly improved, i.e. the means of the ground position differences of over 99% image-pairs are less than one ground sample distance of WFV image (16m). The RMSE is also improved when the geometric distortion is corrected. Several screenshots of the generated DOMs are shown in figure 9. The screenshots show that the mosaicking errors in these areas are less than 2 pixels. It is also revealed by the RMSE in figure 8, that most RMSE of image-pairs are less than 2 GSDs (32m).



Figure 9. Screenshots of Tibet DOMs under condition 2.

#### 4.2 Producing DOMs covering mainland of China

To fully test the automatic processing system, 426 WFV SC images covering the most part of mainland of China were used for trial production. The testing desktop computer has an Intel Core-i7 (four cores and eight threads) CPU, 32GB internal storage, and a Raid-0 disk array constructed with four solid state disks. The whole time cost of the automatic producing is about 2 days (the radiometric correction and block adjustment is not processed). The detailed time costs are listed in Table 3.

works	time cost
cloud detection (automatic)	2 hours
GCPs matching (automatic)	16 hours
distortion correction (automatic)	2 hours
Check the accuracy (manual)	8 hours
DOM generation (automatic)	6 hours

Table 3. Time costs of the processing workflow

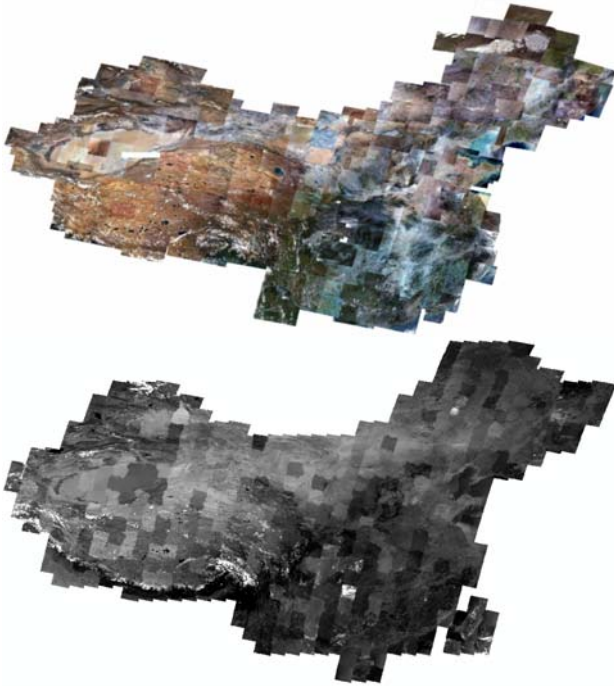


Figure 10. The ortho-rectified GF-1 WFV images (above) and the referencing Landsat-8 images (below).

The accuracy check took our operator about 8 hours and this work can only be processed manually. The DOM generation is executed after the accuracy check to avoid doing this time-costing work again. The land-covers of the WFV images and the referencing Landsat-8 images are illustrated in figure 10.

## 5. CONCLUSION

Having four days revisit time, the GF-1 WFV imaging system enables people to quickly obtain remote sensing images in a large scale. This property enables people to make periodically land-monitoring in many fields, such as crop condition monitoring and management, air pollution monitoring, and land-cover monitoring. However, the improved ability of data-acquisition also promotes the requirements of the degree of automation in data processing. The automatic workflow described in this paper gives a feasible solution that can be universally used in the production of DOM over national wide scale.

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